

# **Space Science : Atmosphere**

## **Greenhouse Effect**

### **Part-5a**

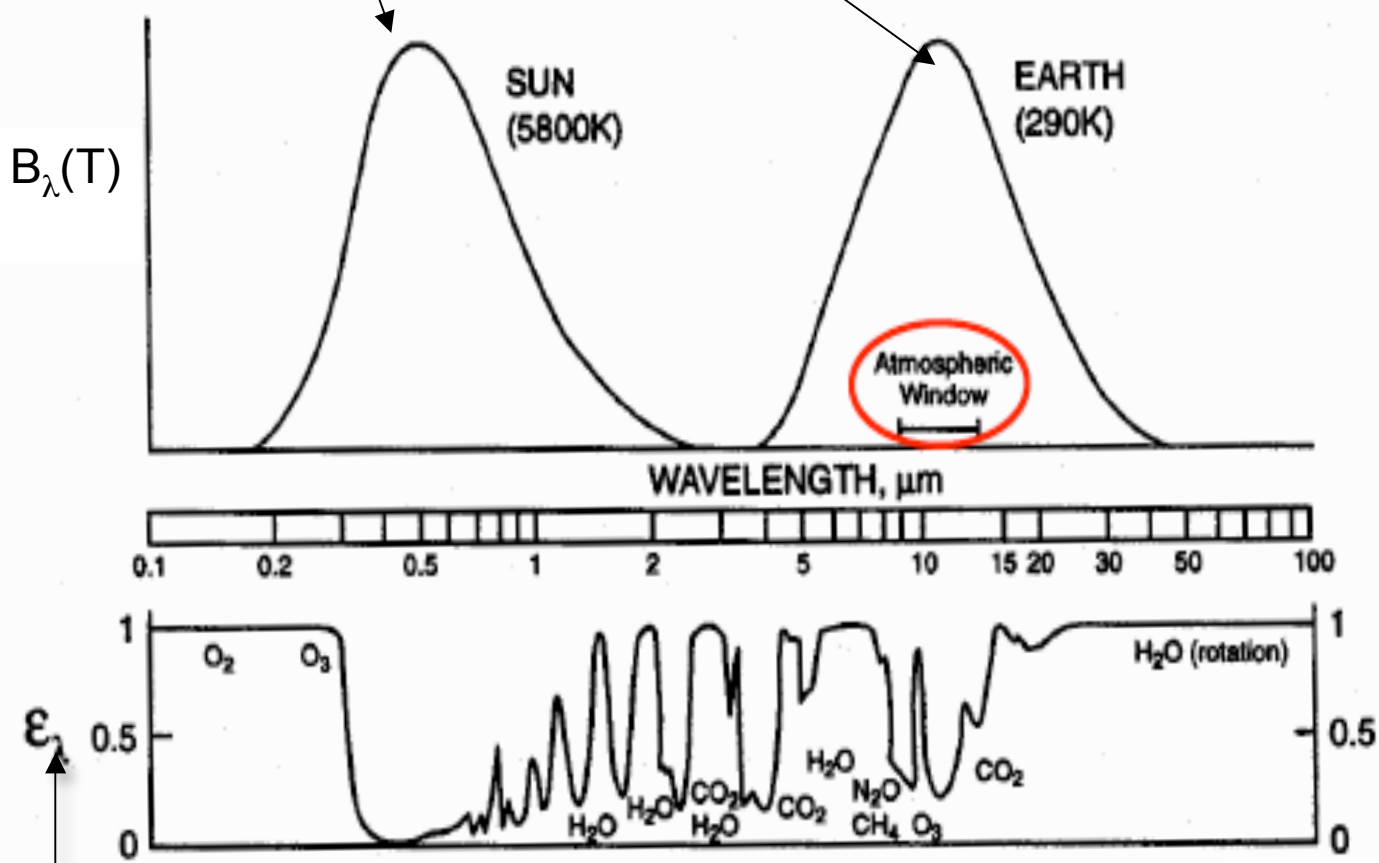
**Solar + Earth Spectrum**  
**IR Absorbers**  
**Grey Atmosphere**  
**Greenhouse Effect**

# Radiation: Solar and Earth Surface

$B_\lambda(T)$  Planck Ideal Emission

Integrate over wave length and angles  
to get total emission from a surface

$$F_{IR} = \sigma T^4 ; \text{Stephan Boltzmann}$$



Fraction absorbed

**Atmosphere is mostly transparent  
in visible but opaque in UV and IR;  
IR window 8-13 $\mu\text{m}$**

# GRAY ATMOSPHERE

(Transparent vs. Gray)

Partially absorbing in IR but absorption is independent of frequency over the range of relevant frequencies

## Process

**Surface heated by visible**

**Warm surface emits IR**

~ 3 – 100  $\mu\text{m}$

peak ~ 15  $\mu\text{m}$

IR absorbed by

CO<sub>2</sub>, O<sub>3</sub>, H<sub>2</sub>O, etc.

**Remember why not O<sub>2</sub> and N<sub>2</sub> ?**

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# Vibrational Bands



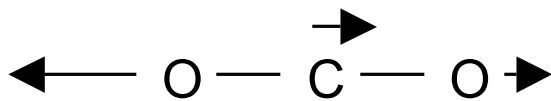
(IR active?)

Symmetric Stretch



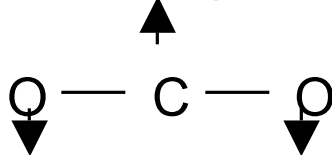
7.46 μm (N)

Asymmetric Stretch

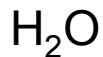


4.26 μm (Y)

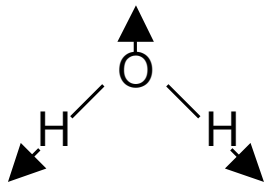
Bending



**15.0 μm (Y)**  
**Near IR peak**

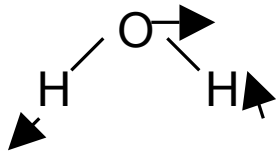


Symmetric Stretch



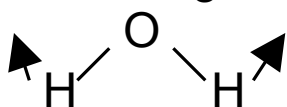
2.73 μm (Y)

Asymmetric Stretch



2.66 μm (Y)

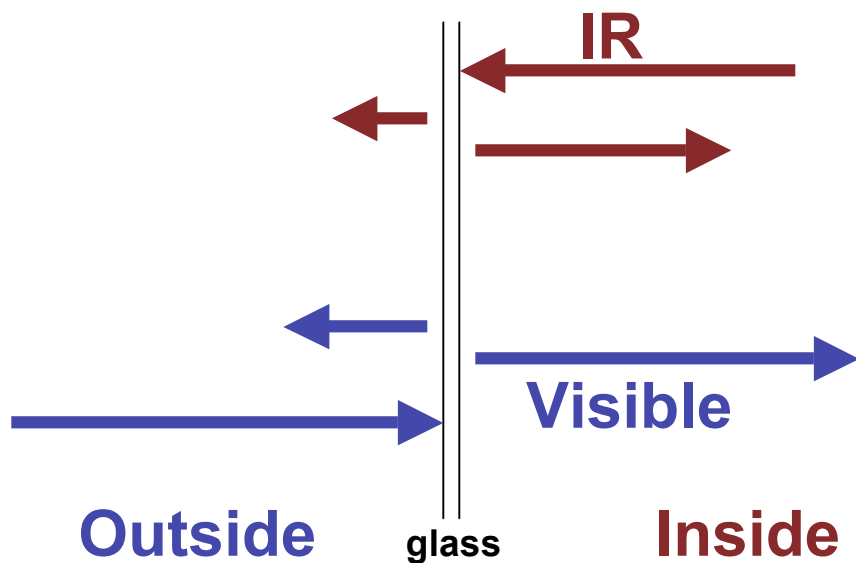
Bending



6.27 μm (Y)

You can have  
combination bands  
or 2 vib. levels

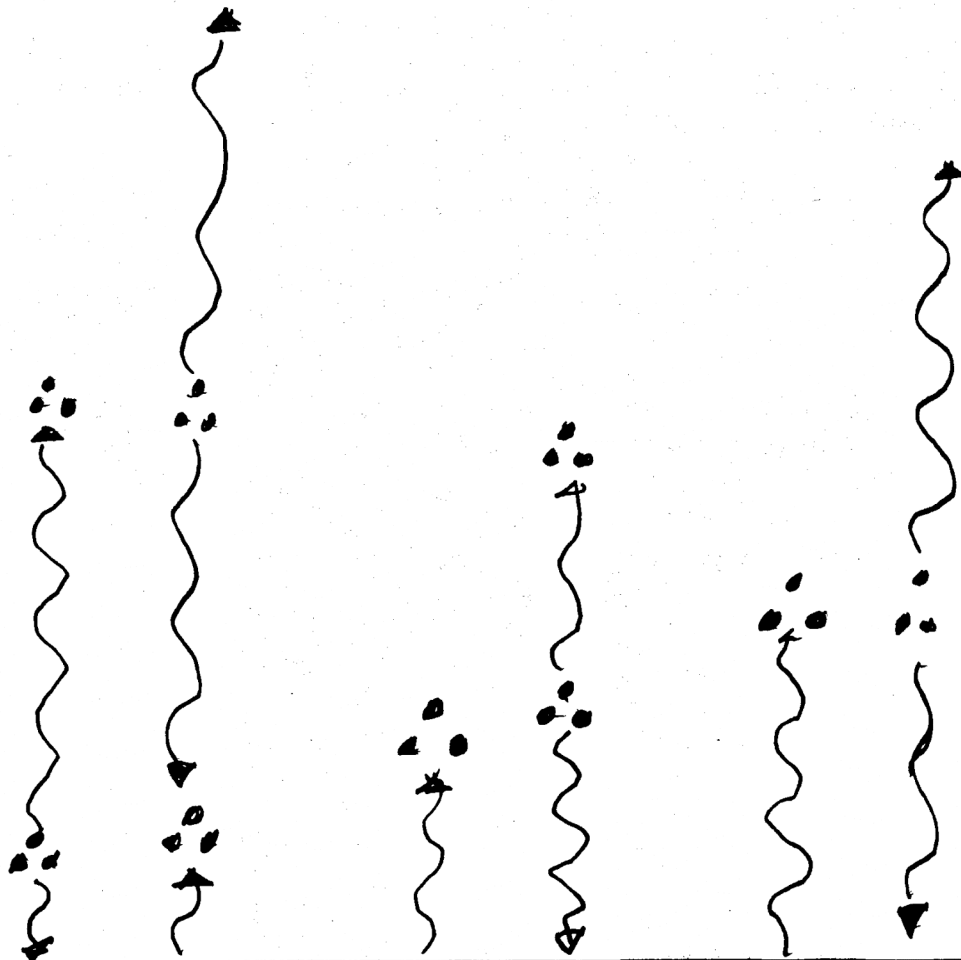
# A Real Green House



How do you get IR out equal to  
Visible light absorbed inside:  
RAISE  $T$

Note: For a real green house  
convection may be as important:  
i.e. glass a thermal barrier

# IR Emission and Absorption



**Ground Emits**

**Primarily triatomic molecules absorb and re-emit: vibrational and rotational states**

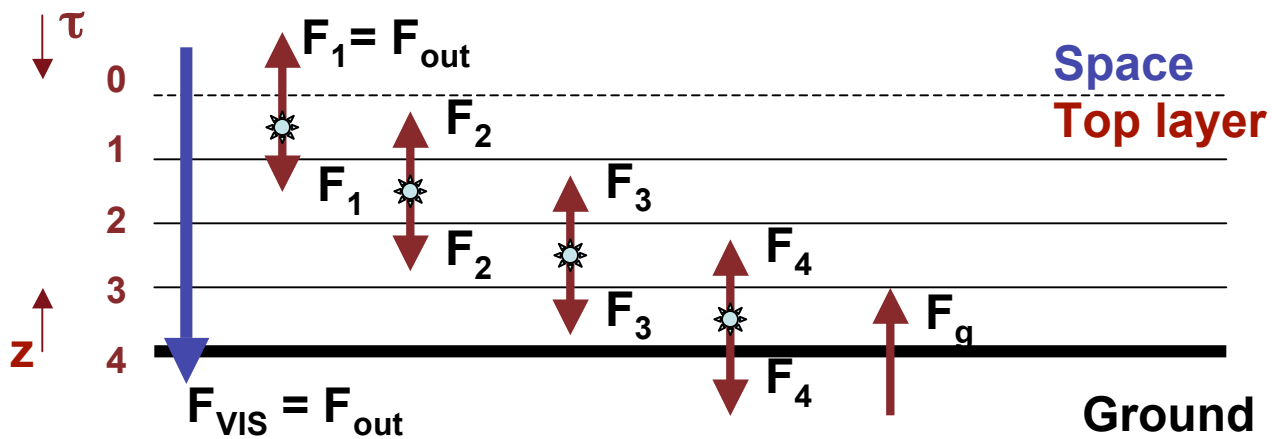
**To determine T we assume excited molecules heat locally by collisions.**



**Also at a given T (CO<sub>2</sub> excited and can emit)**



# Goody + Walker (Simple Model)



I chose an atmosphere 4 layers thick in the IR

Each layer is  $\tau_{\text{IR}} = 1$  thick

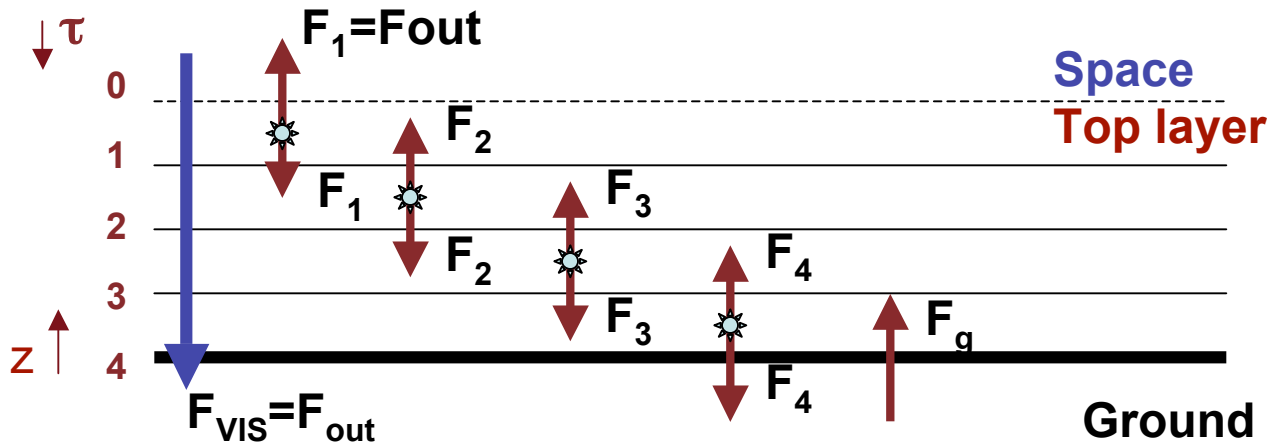
It would absorb  $(1 - e^{-1}) \sim 63\%$  of the flux  
from the layer above and below  
(we'll assume 100% for simplicity)

We will also assume that each layer  
emits like a perfect emitter:  
 $\sigma T^4$ , where  $T$  is its temperature

The surface is heated by the sun and it emits

**Note: heat balance**  $F_{\text{vis}} = F_{\text{out}} = \sigma T_e^4$

# (simple Greenhouse cont.)



Divide atmosphere into IR thick layers:  $\tau_{\text{IR}} \sim 1$

*Energy balance in each layer*

$$0 \quad F_1 = F_{\text{out}} = \sigma T_e^4$$

**Out = In**

$$1 \quad 2 F_1 = F_2 \quad \rightarrow \quad F_2 = 2 F_1 = 2 F_{\text{out}}$$

$$2 \quad 2 F_2 = F_1 + F_3 \quad \rightarrow \quad F_3 = 2 F_2 - F_1 = 3 F_{\text{out}}$$

$$3 \quad 2 F_3 = F_2 + F_4 \quad \rightarrow \quad F_4 = 2 F_3 - F_2 = 4 F_{\text{out}}$$

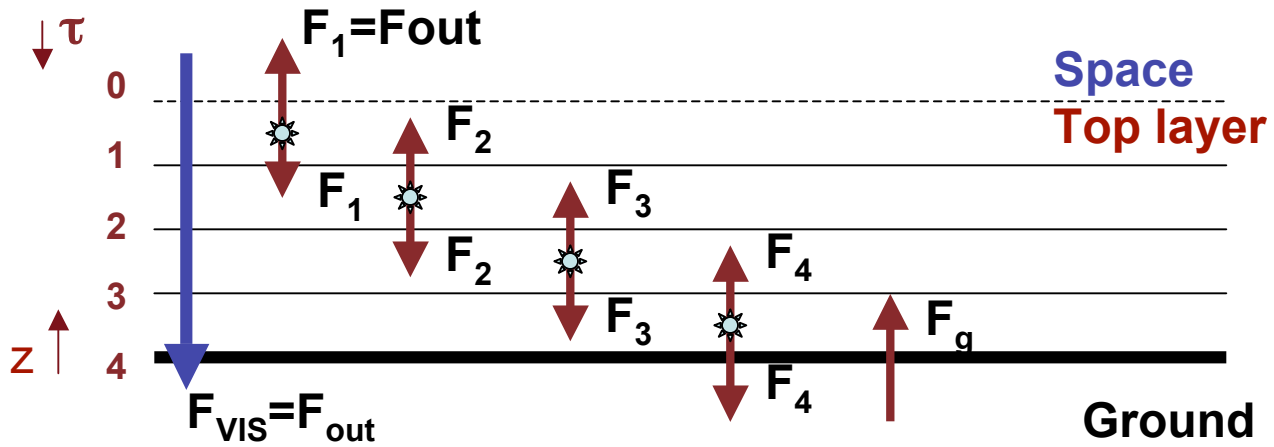
$$4 \quad 2 F_4 = F_3 + F_g \quad \rightarrow \quad F_g = 2 F_4 - F_3 = 5 F_{\text{out}}$$

**Note:** can write  $F_g = [1 + \text{number of layers}] F_{\text{out}}$

$$\sigma T_g^4 = \sigma T_e^4 [1 + \tau_g]$$

$\tau_g$  = number of IR layers

# (simple Greenhouse cont.)



Divided atmosphere into IR layers:  $\tau_{\text{IR}} \sim 1$

We find :  $F_g = [1 + \text{number of layers}] F_{\text{out}}$

$$\sigma T_g^4 = \sigma T_e^4 [1 + \tau_g]$$

**Ground layer (gives the same result)**

$$g \quad F_{\text{vis}} + F_4 = F_g \rightarrow F_g = F_{\text{viz}} + 5 F_{\text{out}}$$

Since  $F_{\text{viz}} = F_{\text{out}} = \sigma T_e^4$ , then get the same result

$$\therefore \text{ drop } \sigma : \quad \underline{\underline{T_g^4 = T_e^4 [1 + \tau_g]}}$$

# How Many Layers on Earth

$$\tau_g \approx 2$$

Using  $T_e = 250\text{K}$

Implies

$$T_g = 330\text{K} ! \quad \text{No Way}$$

Again - it is clear that at the surface  
convection is important  
but it gives  $\sim 288\text{K}$  too cold

**Greenhouse + Convection determine surface T**

(Note  $F_4$  - -air above surface

we have to do a little more carefully)

**Before Finishing: can make an estimate**

$$\text{If } \tau_g \approx 2 \approx \sigma_{\text{abs}} f_{\text{abs}} N$$

Using  $N = 2 \times 10^{25} \text{ mol/cm}^2$  (see early lecture)

$$f_{\text{abs}} = 1\% \quad (\text{H}_2\text{O}, \text{CO}_2, \text{O}_3 \dots)$$

$$\sigma_{\text{abs}} \approx 10^{-23} \text{ cm}^2$$

IR absorbers have small cross sections relative to UV

## **VENUS (Problem for set 3)**

$$T_e = 232 \quad T_g = 750$$

**Therefore:  $\tau_g^* = ?$**

**Then: Use cross section  
from previous slide, pure CO<sub>2</sub>**

**Find column:  $N_{\text{CO}_2} = ?$**

# GREEN HOUSE EFFECT

H <sub>2</sub> O	+ 21 K
CO <sub>2</sub>	+ 7 K
O <sub>3</sub>	+ 2 K
Other(N <sub>2</sub> O, CH <sub>4</sub> , ...)	<u>+ 3 K</u>
	33 K

$$T_e + \Delta T$$

$$255 \text{ K} + 33 \text{ K} = 288 \text{ K}$$

Note: Increase O<sub>3</sub>  $\Rightarrow$   $z > 30\text{K}$  increase T  
 $z < 25\text{K}$  decrease T !

Double CO<sub>2</sub>? (~30% increase since 1800)

$$\Delta T = 2.5 - 5 \text{ K} ! \text{ ATMOSPHERE ONLY}$$

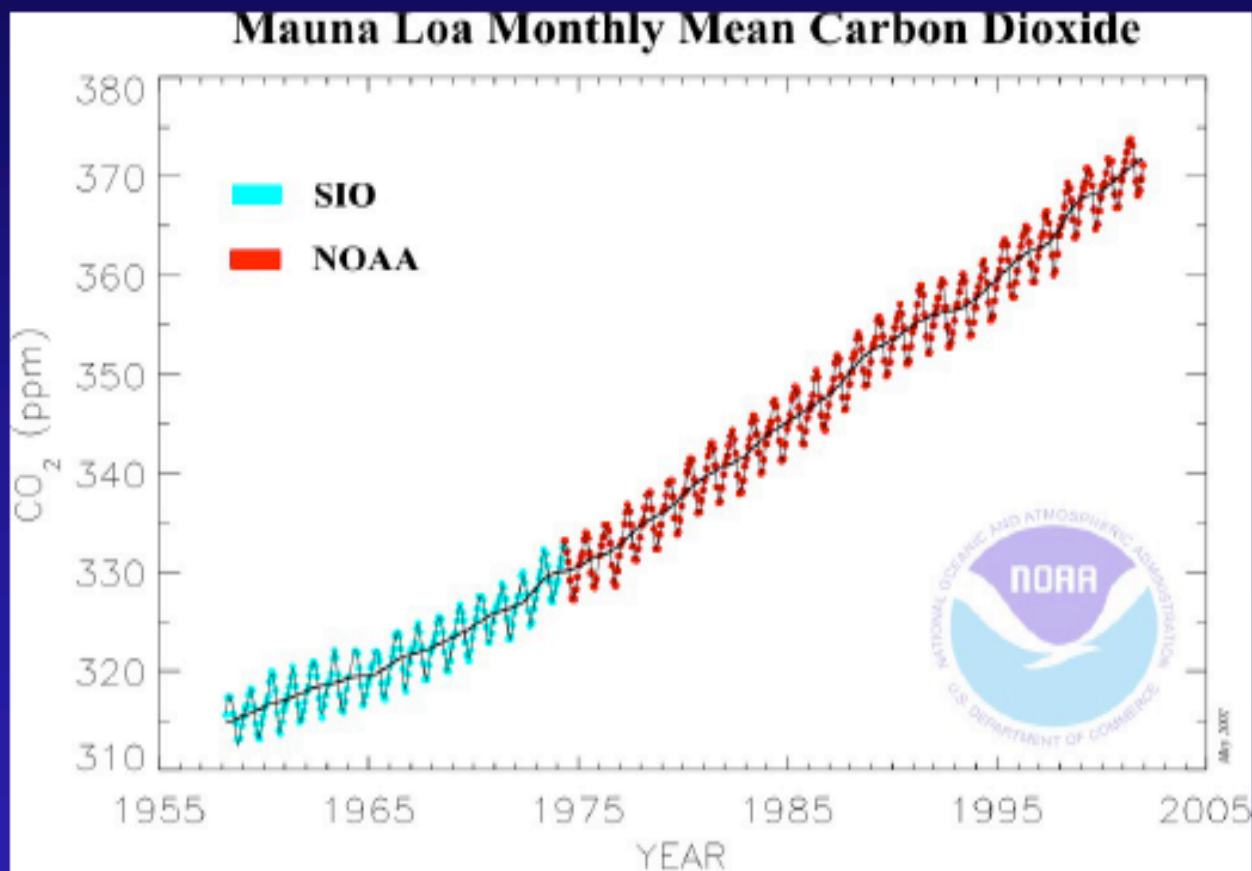
$$= 1 - 2 \text{ K} ! \text{ ATMOSPHERE + OCEAN}$$

We will come back when we discuss atmospheric evolution  
to discuss : Reservoirs for carbon and Time Constants

# Carbon concentration vs. time

- Increase from 280 ppm (1860) to 370 ppm (present) – 31% increase

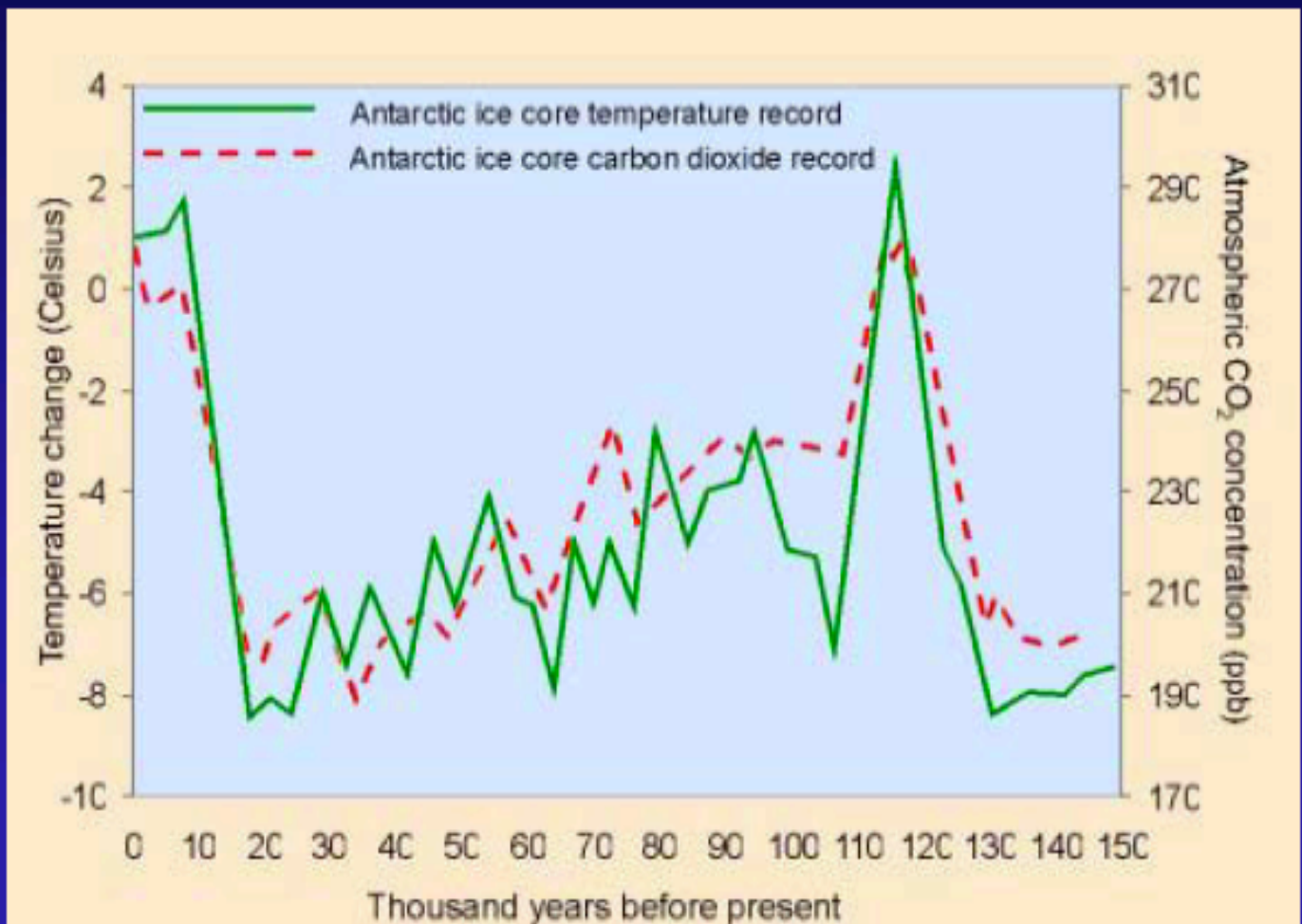
◆ Note: 370 ppm = 0.0370%



Atmospheric carbon dioxide monthly mean mixing ratios. Data prior to May 1974 are from the Scripps Institution of Oceanography (SIO, blue), data since May 1974 are from the National Oceanic and Atmospheric Administration (NOAA, red). A long-term trend curve is fitted to the monthly mean values. Principal investigators: Dr. Pieter Tans, NOAA CMDL Carbon Cycle Greenhouse Gases, Boulder, Colorado, (303) 497-6678, [ptans@cmdl.noaa.gov](mailto:ptans@cmdl.noaa.gov), and Dr. Charles D. Keeling, SIO, La Jolla, California, (616) 534-6001, [cdkeeling@ucsd.edu](mailto:cdkeeling@ucsd.edu).

# Carbon Concentration Long Term

- Historical Record – Correlation of CO<sub>2</sub> and Temp



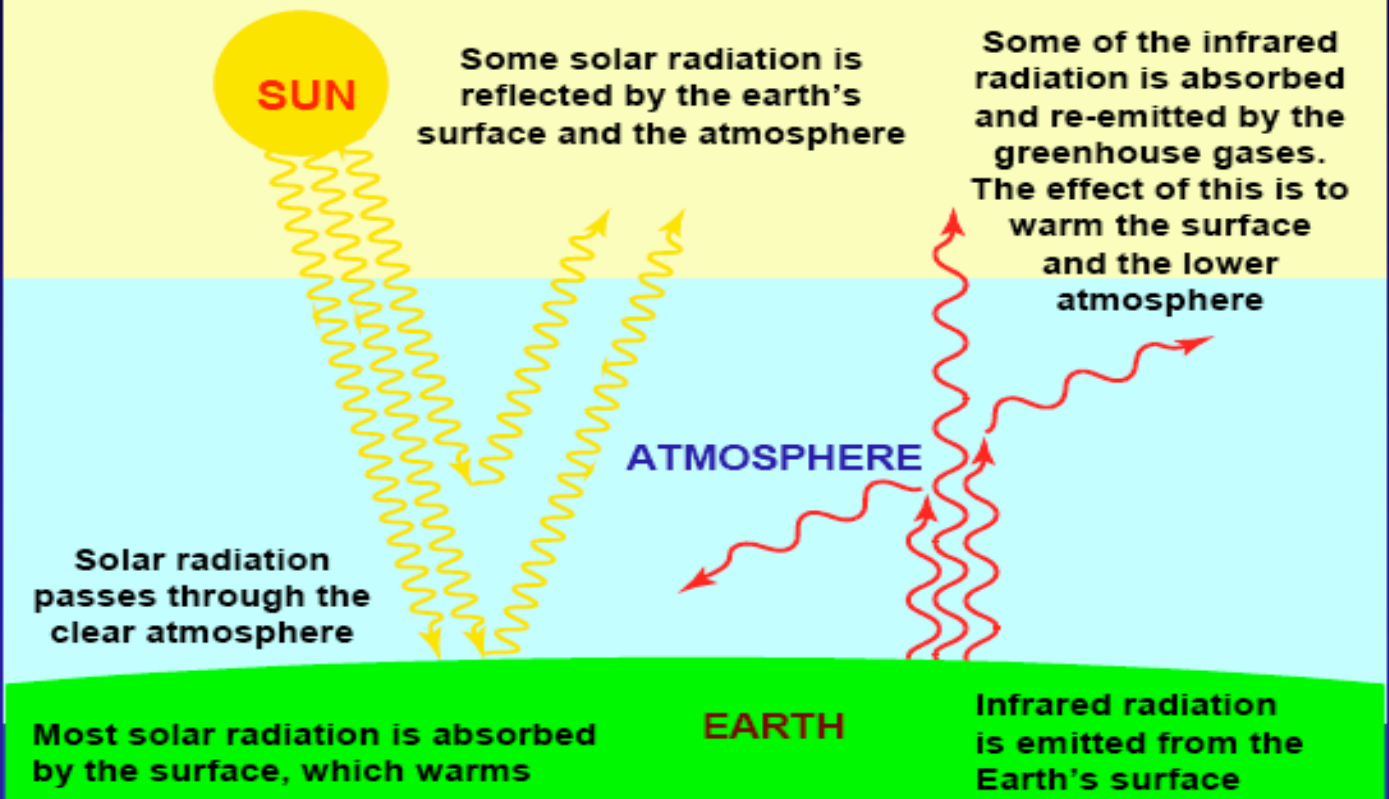
Later we will look at the carbon cycle

- **Carbon Dioxide, CO<sub>2</sub>**
  - ◆ **Contributes 60% of warming**
- **Trace Gases ( 100 –1000 times less)**
  - ◆ **Contribute 40% of warming**
  - ◆ Methane, CH<sub>4</sub> (15-30% of warming)
  - ◆ Nitrogen Oxides, NO<sub>x</sub> (up to 15% of warming)
  - ◆ Chlorofluorocarbons, CFCs (12% of warming)

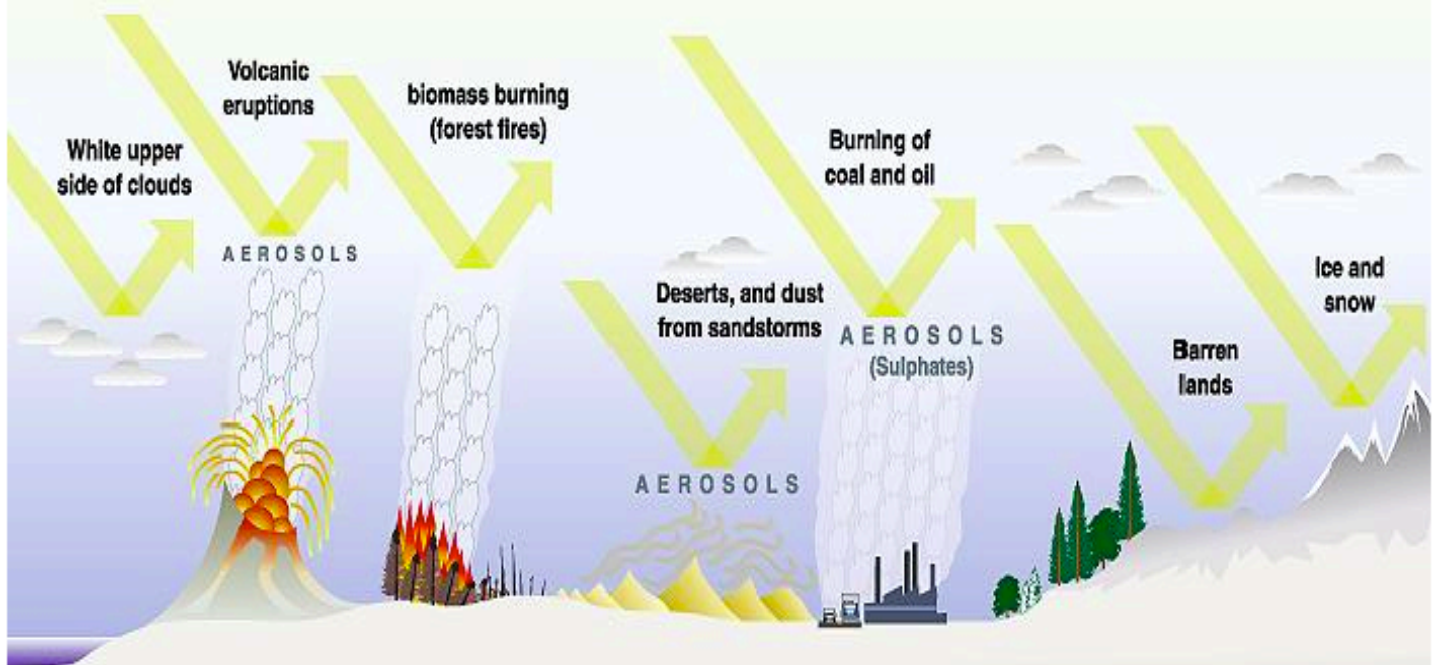
### **However,**

- ❖ Molecule for molecule, CO<sub>2</sub> is less efficient than other greenhouse gases because its atmospheric concentration is high and hence its absorption bands are nearly saturated.
- ❖ Over a 100-year time horizon, reducing SF<sub>6</sub> emissions by 1 kg is as effective from a greenhouse perspective as reducing CO<sub>2</sub> emissions by 24,900 kg.

# Greenhouse Effect is Complex



## The cooling factors



Energy reflected

**Albedo:** ability of a surface to reflect light.

**Aerosols:** tiny particles of liquid or dust suspended in the atmosphere (most important anthropogenic aerosol is sulphate produced from  $\text{SO}_2$ )

**GRID**  
Arendal

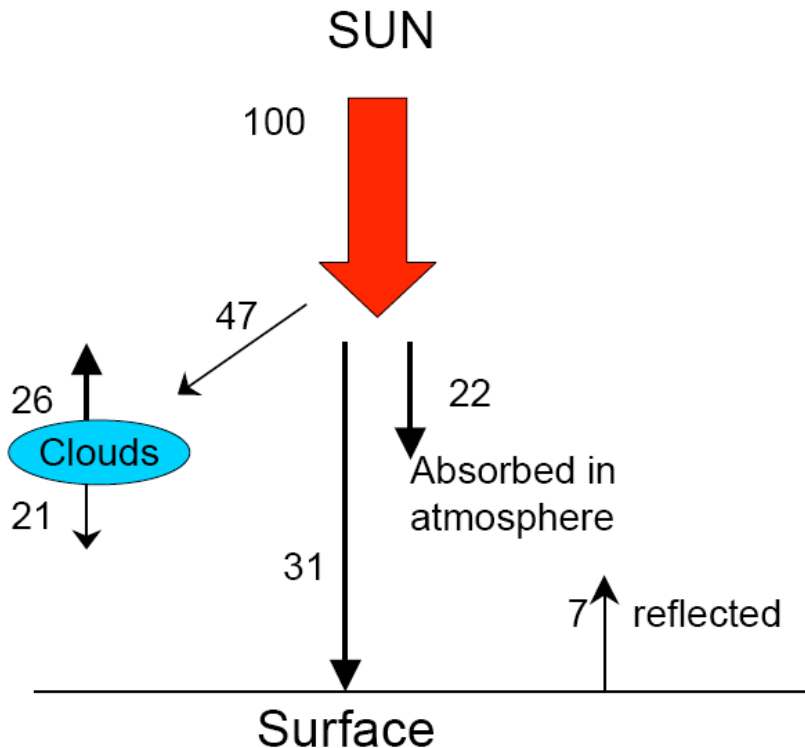
GRAPHIC DESIGN: PHILIPPE REKACEWICZ

# PLANETARY ENERGY BALANCE

G+W fig 3-5

## Incoming solar radiation

At Earth



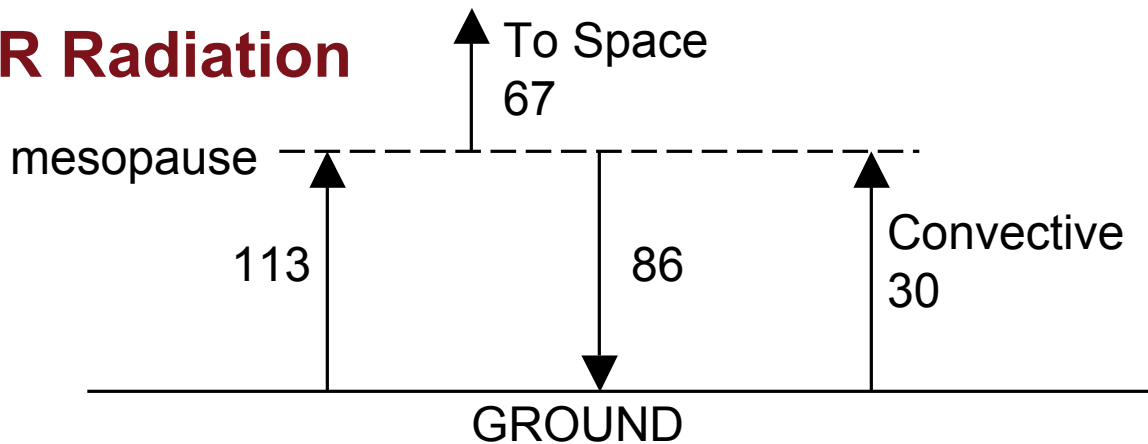
### Absorbed

Clouds	21
Atmos.	22
Ground	<u>24</u>
	67

### Back to space

Reflect	Ground	7
	Clouds	<u>26</u>
<u>Albedo</u>		33

## IR Radiation



## **#5a Summary**

### **Things you should know**

**Planck Radiation Law**

**IR Molecules**

**Greenhouse Effect**

**Surface temperature**